

# New ways to create longer-lived dust asymmetries

- (1) Re-triggered Vortices
- (2) High-mass planets
- (3) Planet Migration
- (4) VSI



## Context

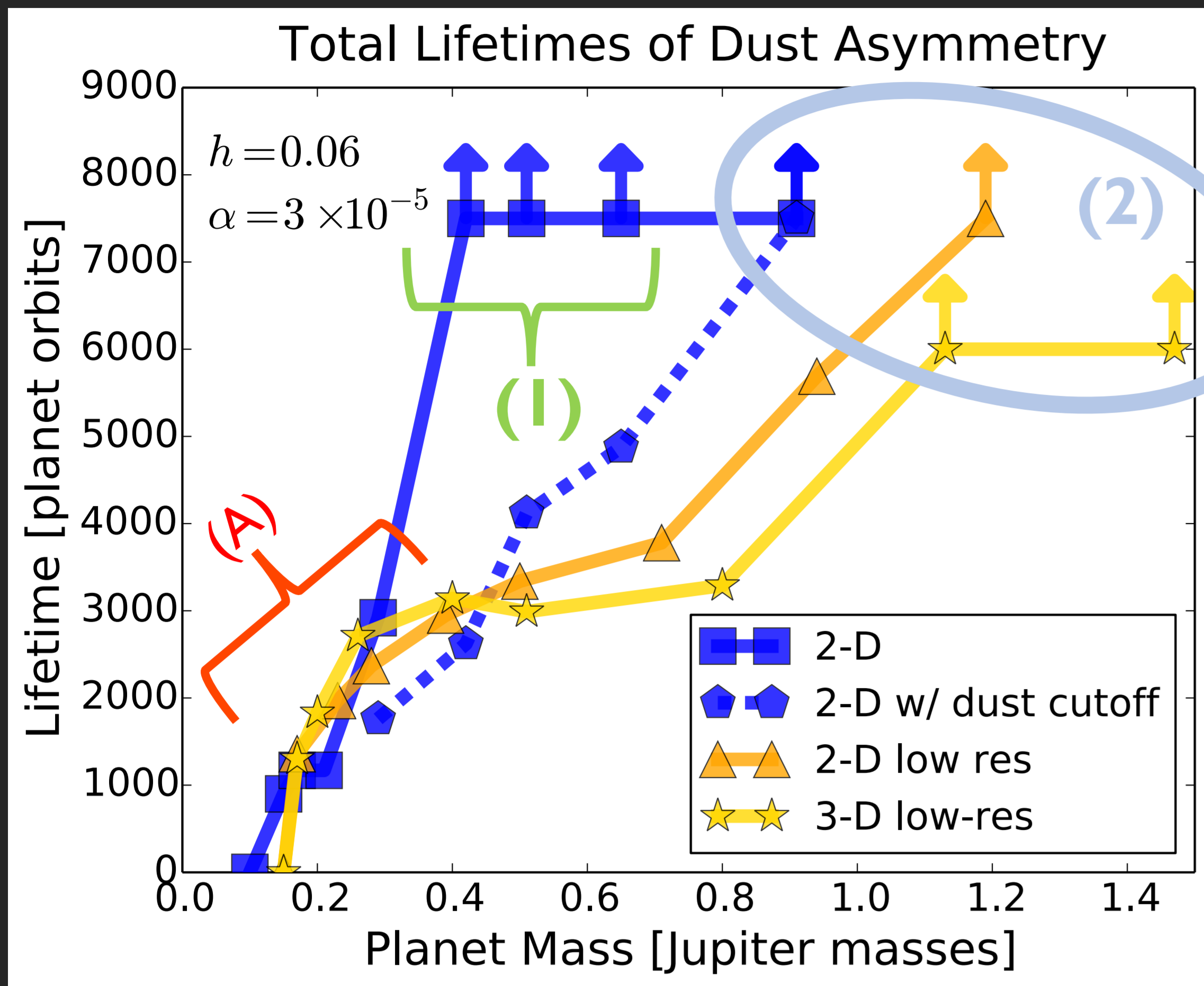
Our previous work (MH+ 17, 19, 21) found that **incorporating the planet's growth** into simulations shortens the lifetimes of a vortex's dust asymmetry.

That's because the resulting vortices are elongated instead of compact.

Only low-mass planets can produce long-lived dust asymmetries by re-triggering vortices (MH+ 21).

## Doesn't work! (MH+ 23)

- (A) Low-mass planets still re-trigger long-lived vortices in high-mass discs, but not in low-mass discs.
- (B) With high enough vertical resolution in 3-D, the dust clumps produced by dust feedback no longer survive indefinitely.

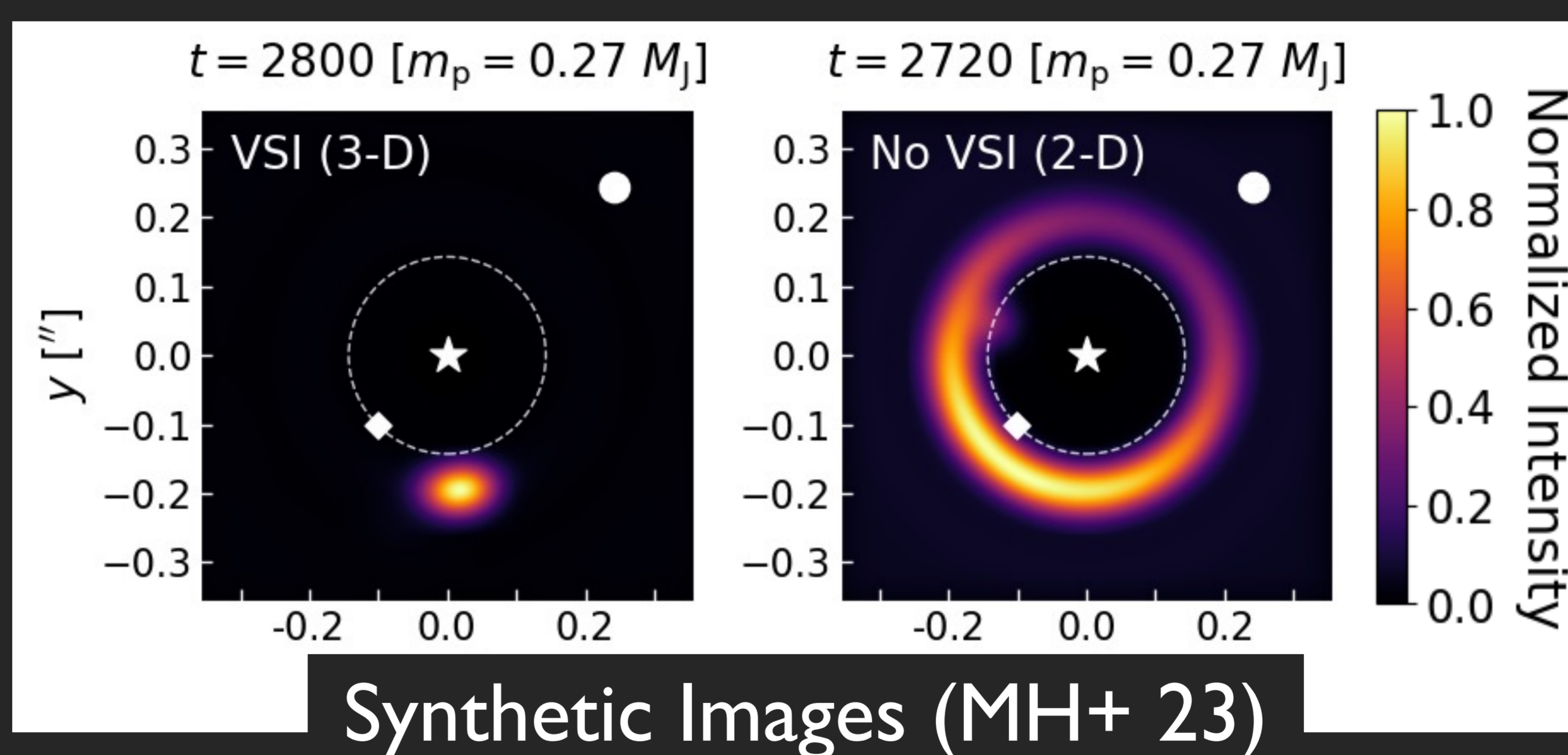


## How it happens!! (MH+ 23)

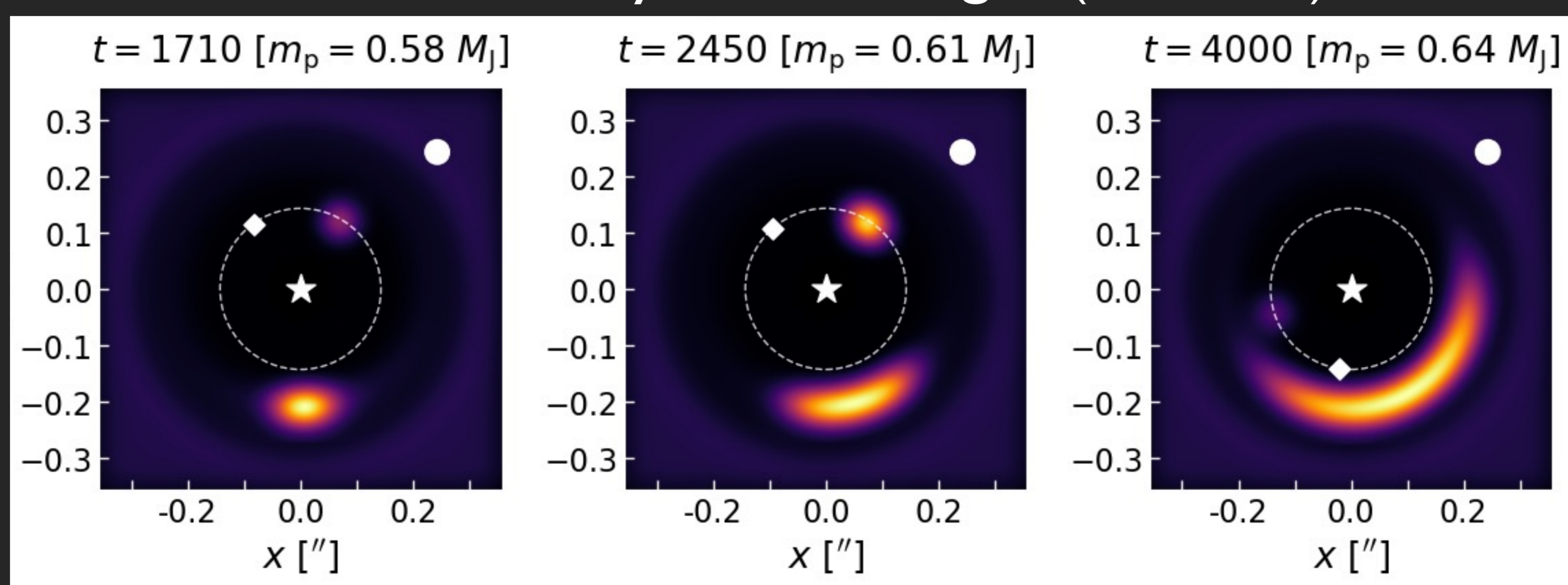
- (1) With high enough resolution ( $29/h$ ), high-mass planets re-trigger vortices early. These new elongated vortices develop **compact cores** that **may not decay in the dust** even after the gas vortex is gone.
- (2) Even though Jupiter-mass planets still induce elongated vortices, they're compact enough that they **do not decay in the gas**.
- (3) The vortex's dust asymmetry can survive indefinitely because it is **left behind** when the planet migrates inwards.
- (4) The VSI can seed the RWI with **compact vortices** that do not decay.

We found four new ways to make **planet-induced vortices survive longer** in simulations, which **contradicts how rare** asymmetries are in mm-dust images of protoplanetary discs.

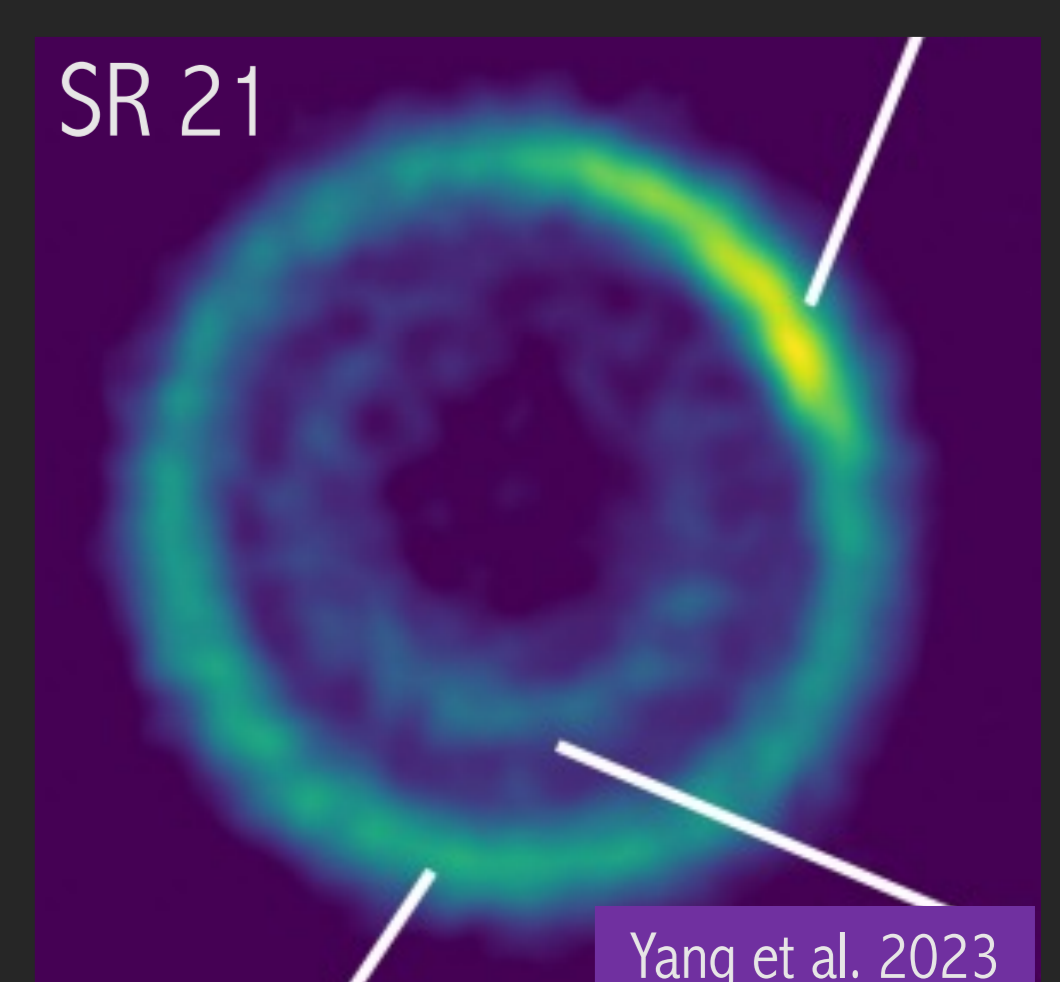
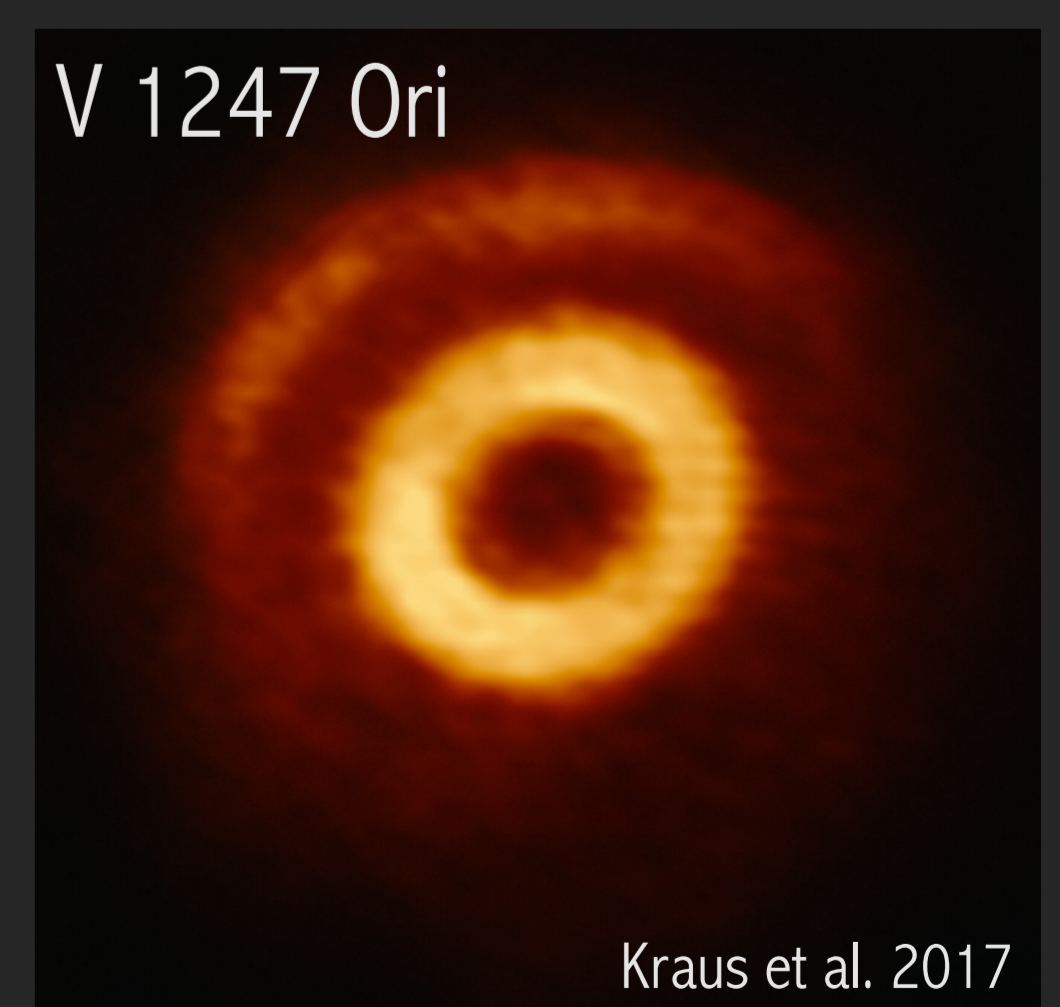
With a low-mass planet, planet-induced vortices are much more **compact with active VSI** than with no VSI.



Once the last **re-triggered vortex** decays in the gas, its compact core slowly decays towards a ring.



## Planet-induced Vortex Candidates (at gap edges)



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Hydro simulations done w/ FARGO3D

MH+ 23 = How to form compact & other longer-lived planet-induced vortices: VSI, planet migration, or re-triggers, but not feedback



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